

# IE350

# Alternate Energy Course

## Lecture # 4

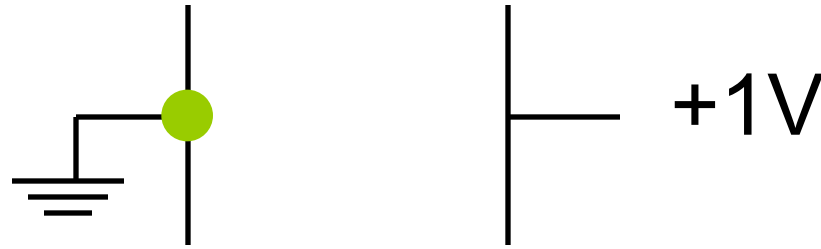
## Energy and Power, Solar Astronomy

# Energy Units - Calorie

- Calorie (cal) = heat to increase by  $1^{\circ}\text{C}$  the 1 gram of water.
- $1 \text{ cal} \approx 4.184 \text{ Joules}$

# Very Small Energy Unit, eV

- Electronvolt (eV) - the amount of kinetic energy gained by a single unbound electron when it passes through an electrostatic potential difference of one volt, in vacuum.



- $1 \text{ eV} = 1.6 \times 10^{-19} \text{ J}$

# Energy unit conversion factors

Solar Energy

	A	B	C	D	E	F	G	H
1	<i>Relationships between different units of measurement for Energy</i>							
2								
3		<b>INPUT</b>	Calory	Joule	Kilowatt-hour	Erg	Electron-volt	kG*m
4	Calories	1	1	4.19E+00	1.16E-06	4.18E+07	2.61E+19	4.27E-01
5	Joules	1.00E+00	2.39E-01	1	2.78E-07	1.00E+07	6.24E+18	1.02E-01
6	Kilowatt-hours	1	8.60E+05	3.60E+06	1	3.60E+13	2.25E+25	3.67E+05
7	Ergs	1	2.39E-08	1.00E-07	2.78E-14	1	6.24E+11	1.02E-08
8	Electron-volts	1	3.83E-20	1.60E-19	4.45E-26	1.60E-12	1	1.63E-20
9	kG*m-s	1	2.34E+00	9.81E+00	2.73E-06	9.81E+07	6.12E+19	1

Storage Calculation / SOLTRM / E for E(PV) / Ph H2 / Fuel / H-nyu Lambda / MultyS WT / En Acc / En Acc (2) / ENERGY\_U / PV costs / PV Pumping

# Energy and Power

- If power is constant

$$\mathbf{E = P \cdot t, \quad P = E/t}$$

- If power is variable and depends on time

$$\mathbf{E = \int P(t)dt, \quad P(t) = dE(t)/dt}$$

# Power Units

- Watt (W) = using one J in one second.
- kW = 1000 W
- Horsepower = 735 W  
= 0.735 kW
- MW = 1000 kW

# Power vs. Energy

- Thus, power is the rate of the energy use.
- Energy is what you pay for repeatedly, as much as you use the energy, the kWh-s – **variable**, operational cost.
- Power is the capacity to use the energy
- You pay for the capacity usually upfront, **fixed** or installation cost.
- E.g. if you decide to buy an air conditioner, you need to solve a power sizing problem. You pay the fixed amount. Later you usually use only a fraction of the total capacity.

# Solar Energy

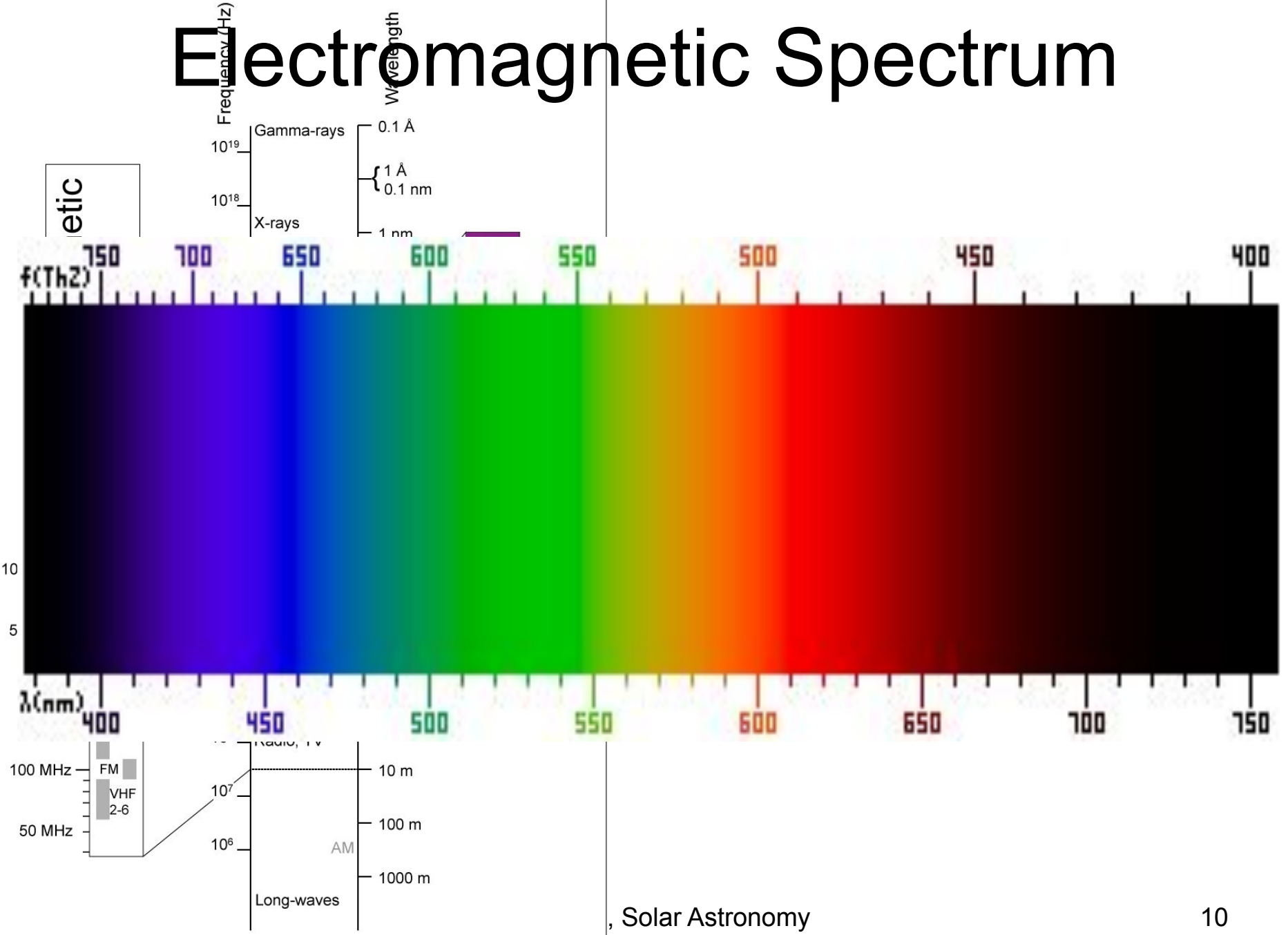
- The SUN:
- Fusion in the sun – the process
- Temperature of the suncrust, black-body radiation – BBR
- Photon energy, light speed, duality
- Electromagnetic Spectrum
- The solar radiation spectrum
- Solar constant =  $1366 \text{ W/m}^2$ .



# The light: particle, wave

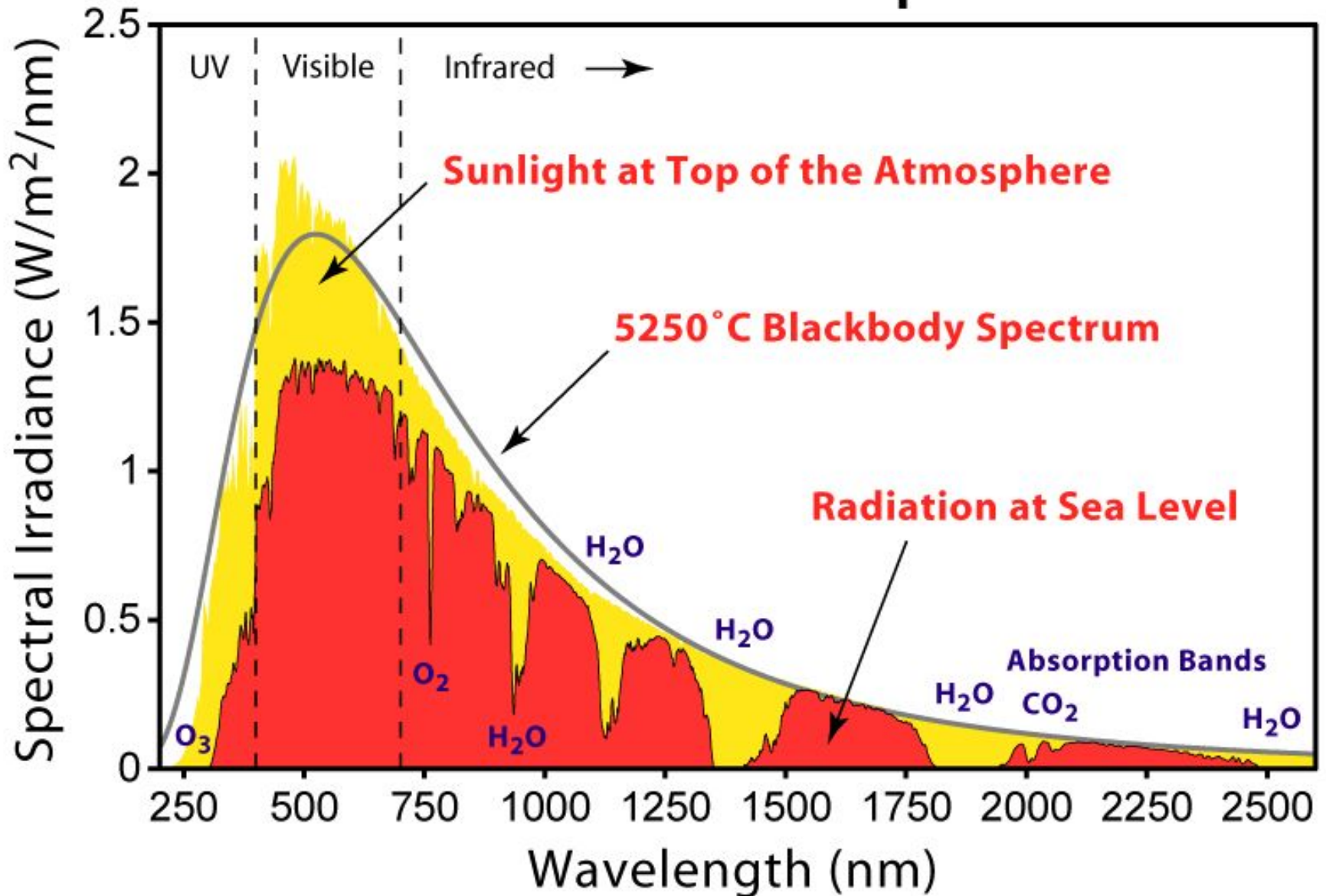
- Particle and wave
- Light speed,  $c = 299,792,458 \text{ m/s}$   
 $c \approx 300,000 \text{ km/s}$
- Photon energy,  $E = h\nu$ ,  $\nu = \text{frequency}$ ,
- $h$  is Planck's constant,  $h = 6.626 \cdot 10^{-34} \text{ J s}$   
 $h = 4.135 \cdot 10^{-15} \text{ eV s}$ .
- $\lambda = c/\nu$
- $E = hc/\lambda$

# Electromagnetic Spectrum



, Solar Astronomy

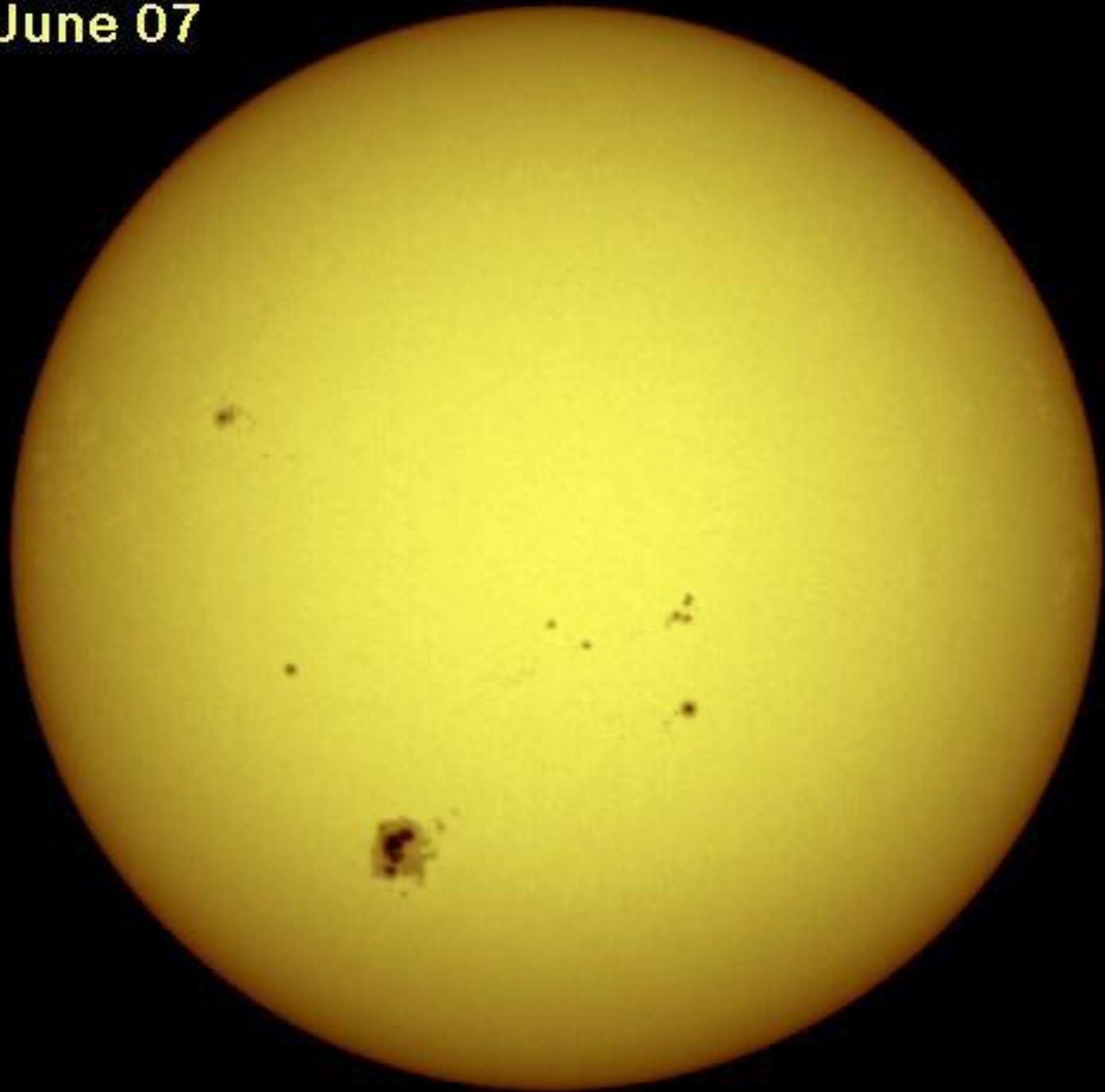
# Solar Radiation Spectrum



# The Sun

- Sun has a capacity of  $3.86 \times 10^{26}$  W  
 $3.86 \times 10^8$  EJ/s
- Earth gets only two-billionth part of it.
- 127,400,000 km<sup>2</sup> - Earth cross-section
- $1.740 \times 10^{17}$  W = 0.174 EJ/s
- Armenian annual energy consumption:  
0.1752 Quads
- Solar Constant = 1366 W/sq.m.
- Average Insolation =  $\frac{1}{4}$  of solar const.  
= 342 W/sq.m.

1992 June 07



# How this energy is generated?

- About 74% of the Sun's mass is hydrogen, 25% is helium, and the rest is made up of trace quantities of heavier elements.

# How this energy is generated?

- The Sun has a surface temperature of approximately 5,500 K, giving it a white color, which, because of atmospheric scattering, appears yellow.

# How this energy is generated?

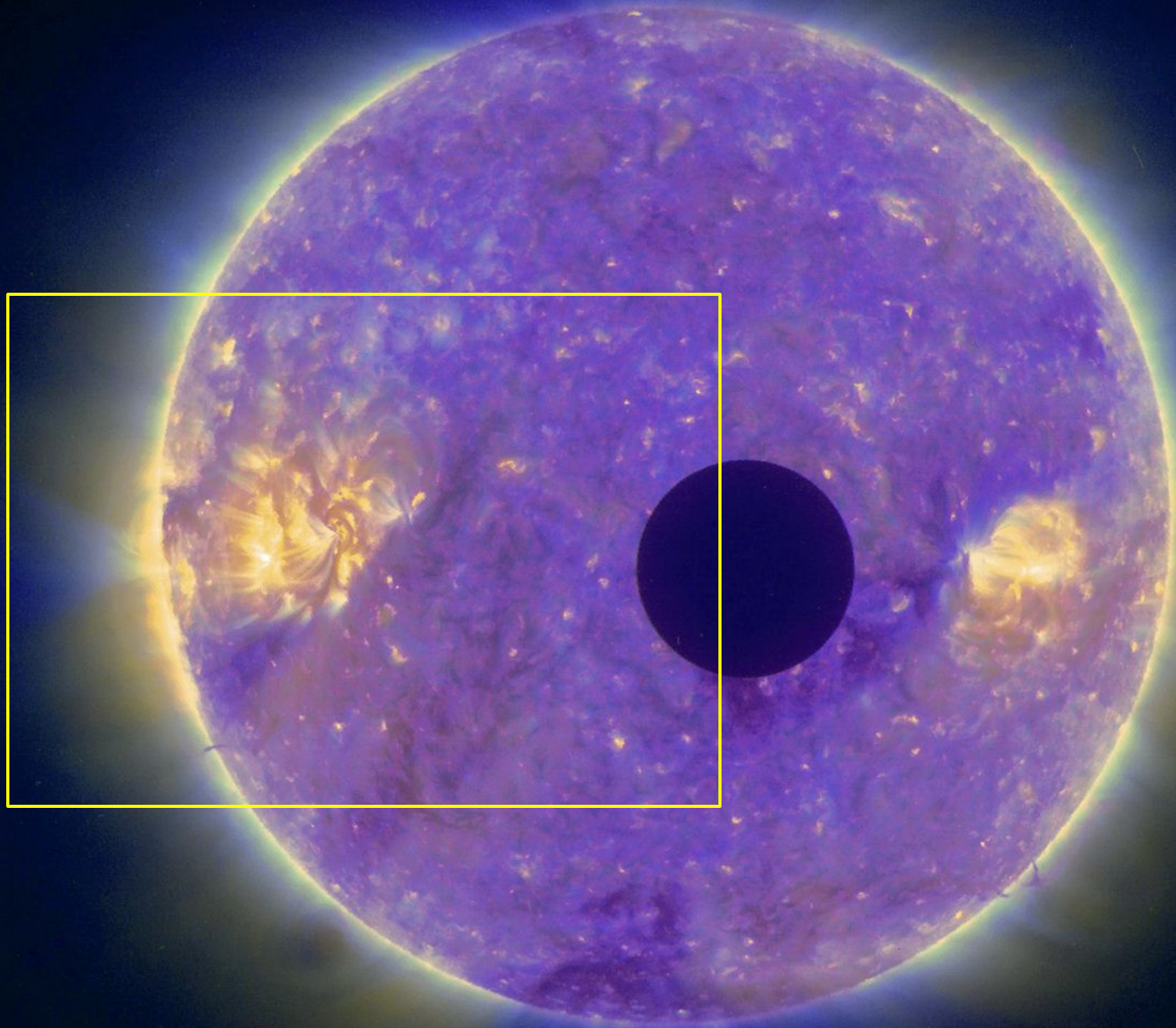
- The Sun diameter:  
 $1.4 \times 10^6$  km = 109 that of the earth.
- Distance from Earth:  
 $1.5 \times 10^8$  km, = 8.31 min at light speed



**How this energy is generated?**

It was *Albert Einstein* who provided the essential clue to the source of the Sun's energy output with his mass-energy relation:

$$\mathbf{E=mc^2}$$



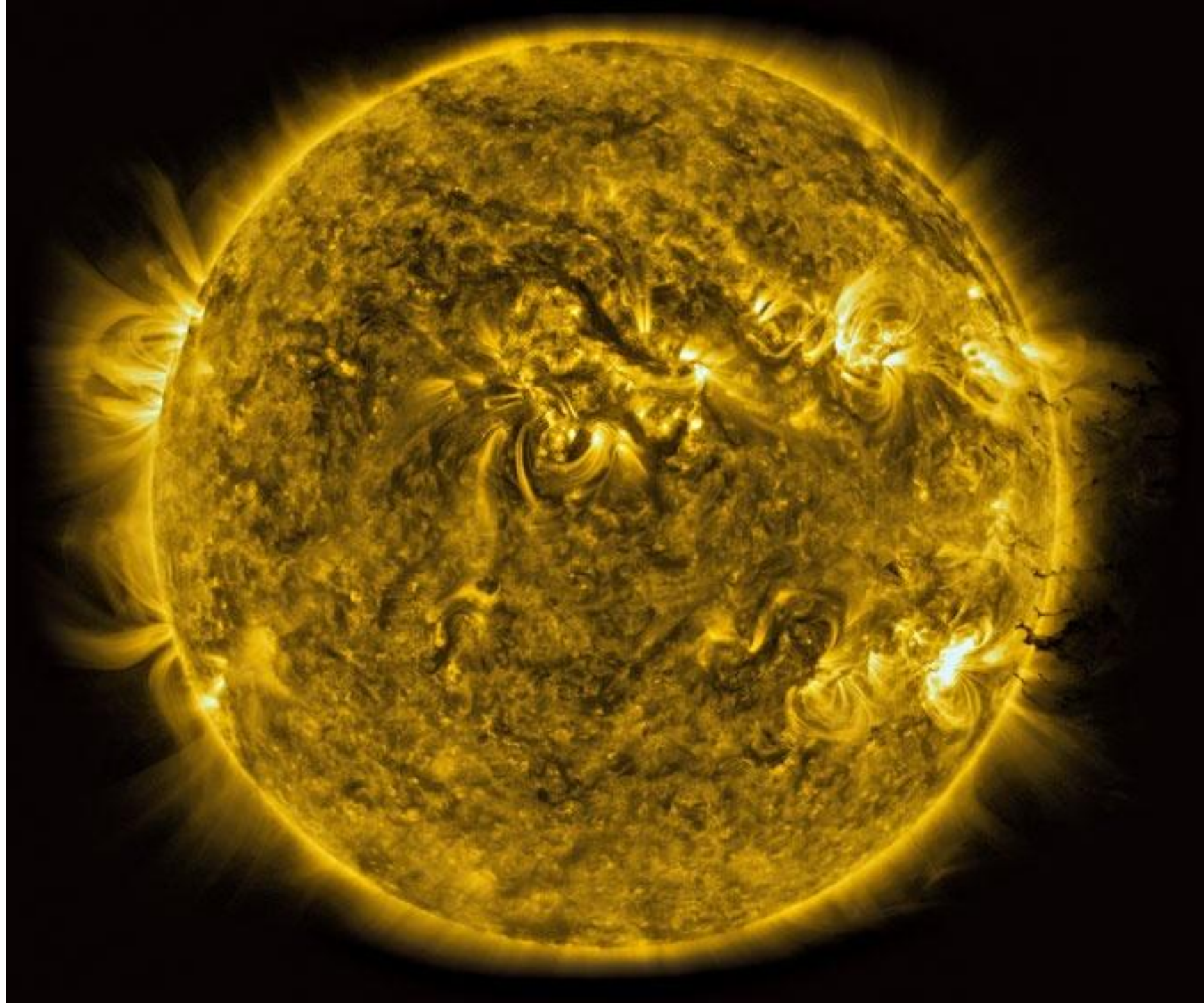


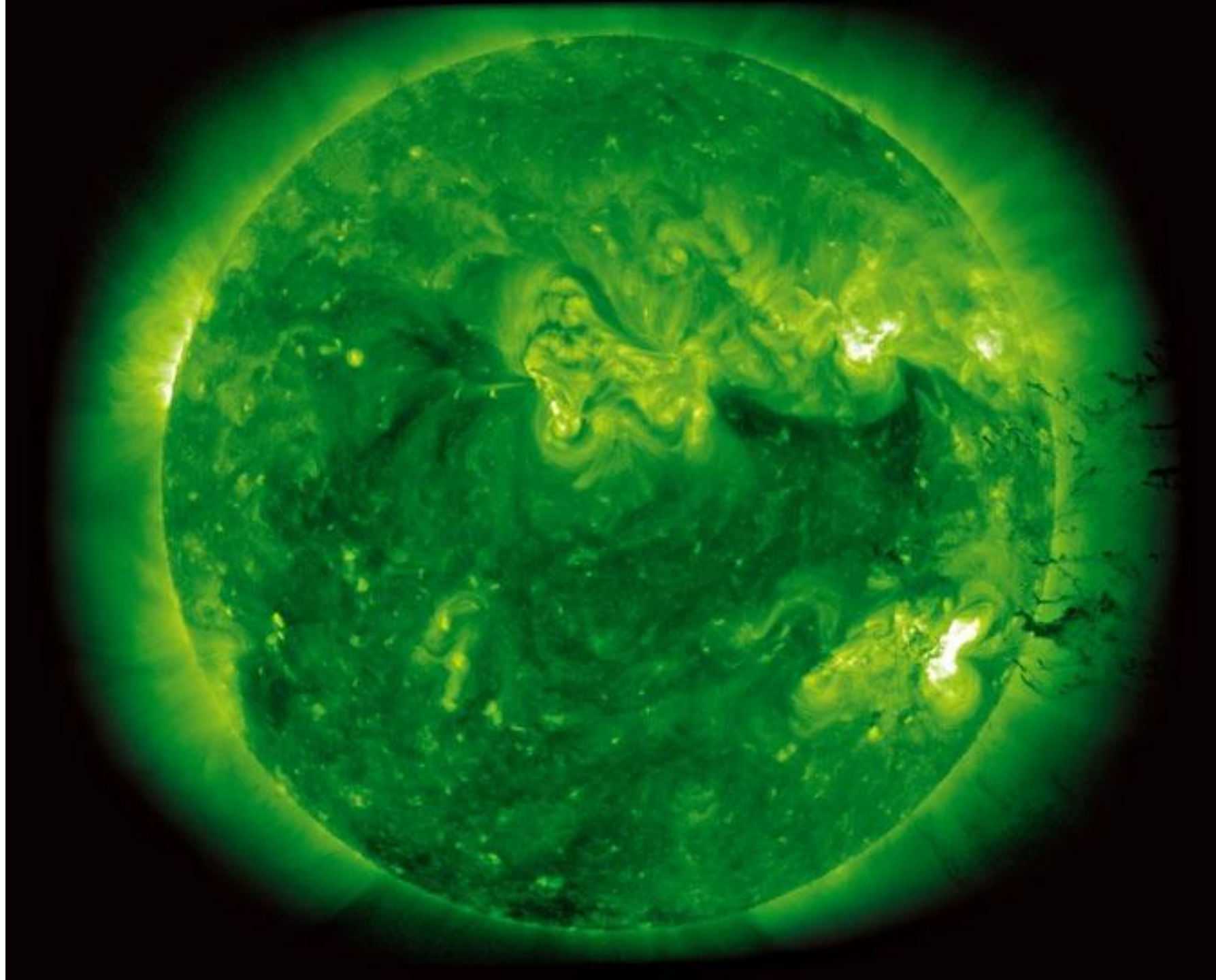


# The Sun

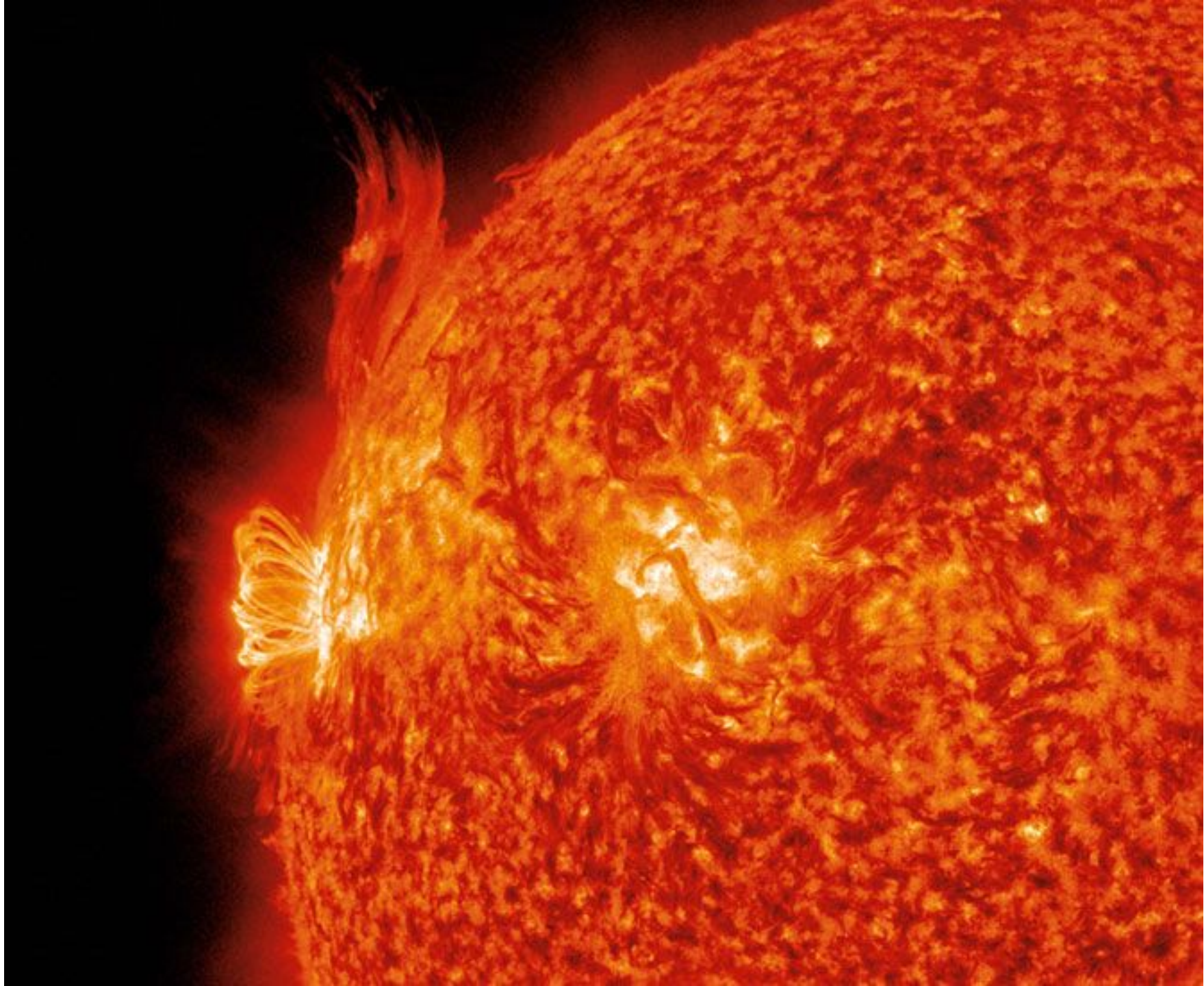


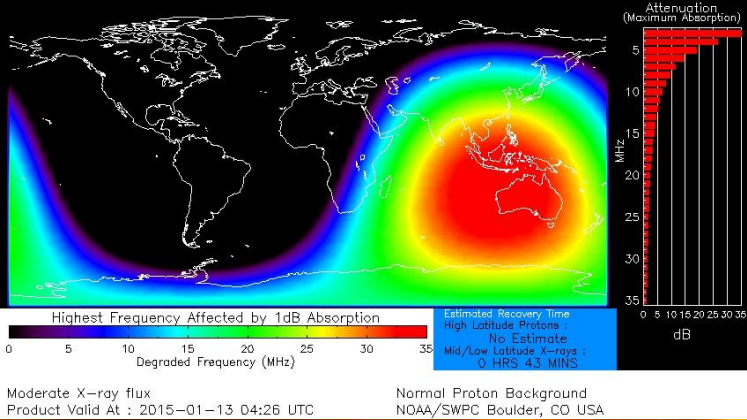






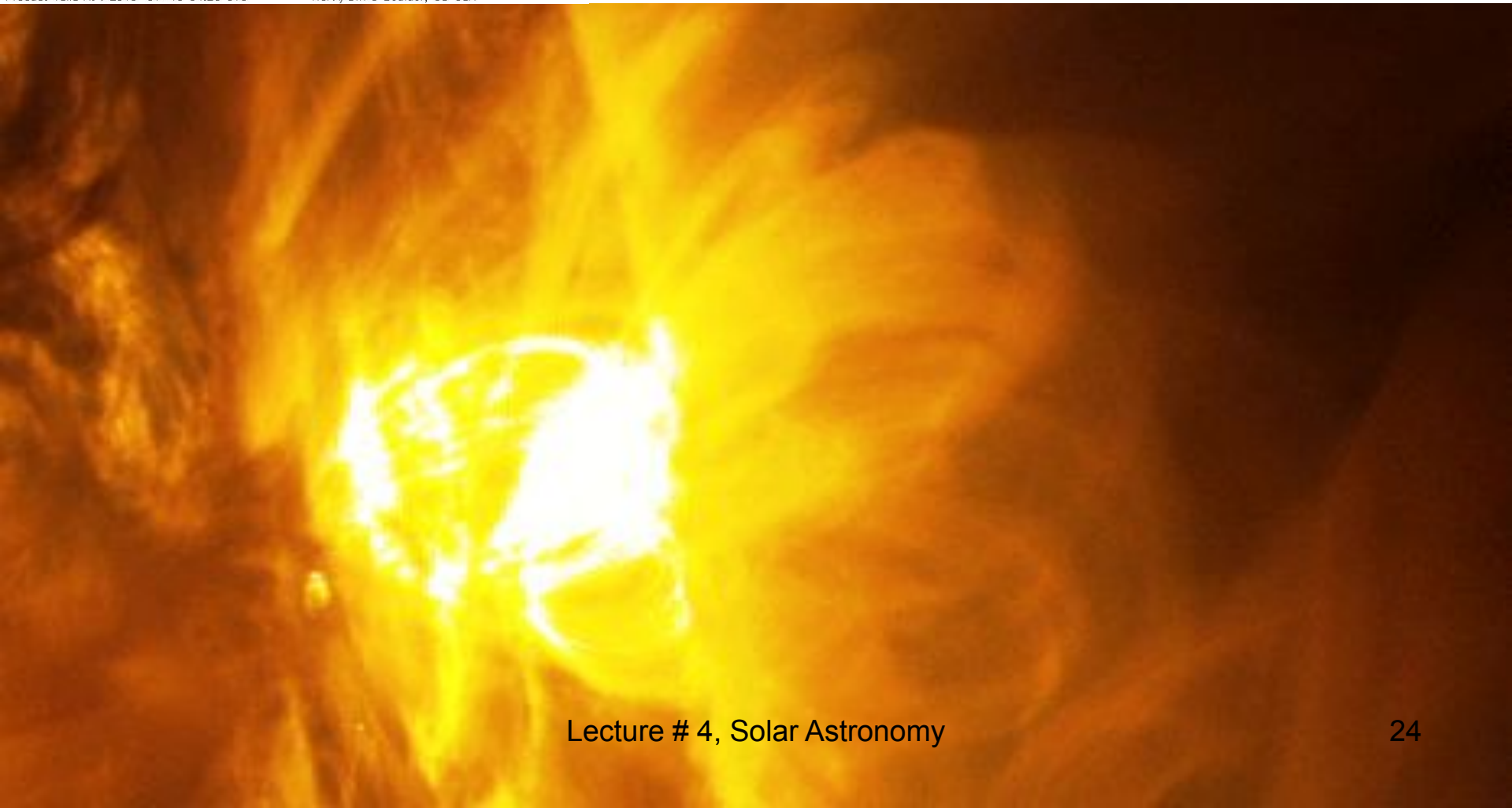




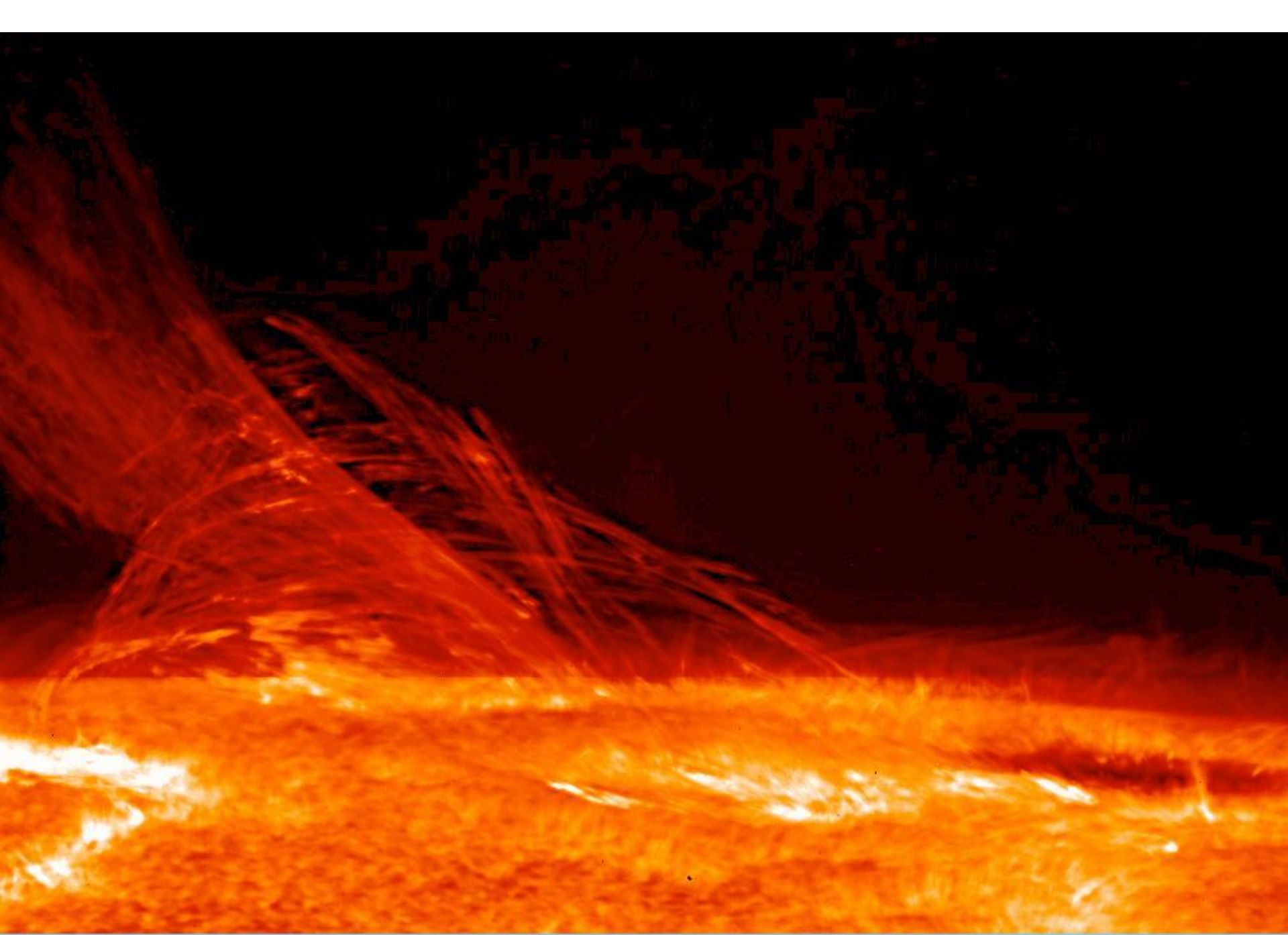


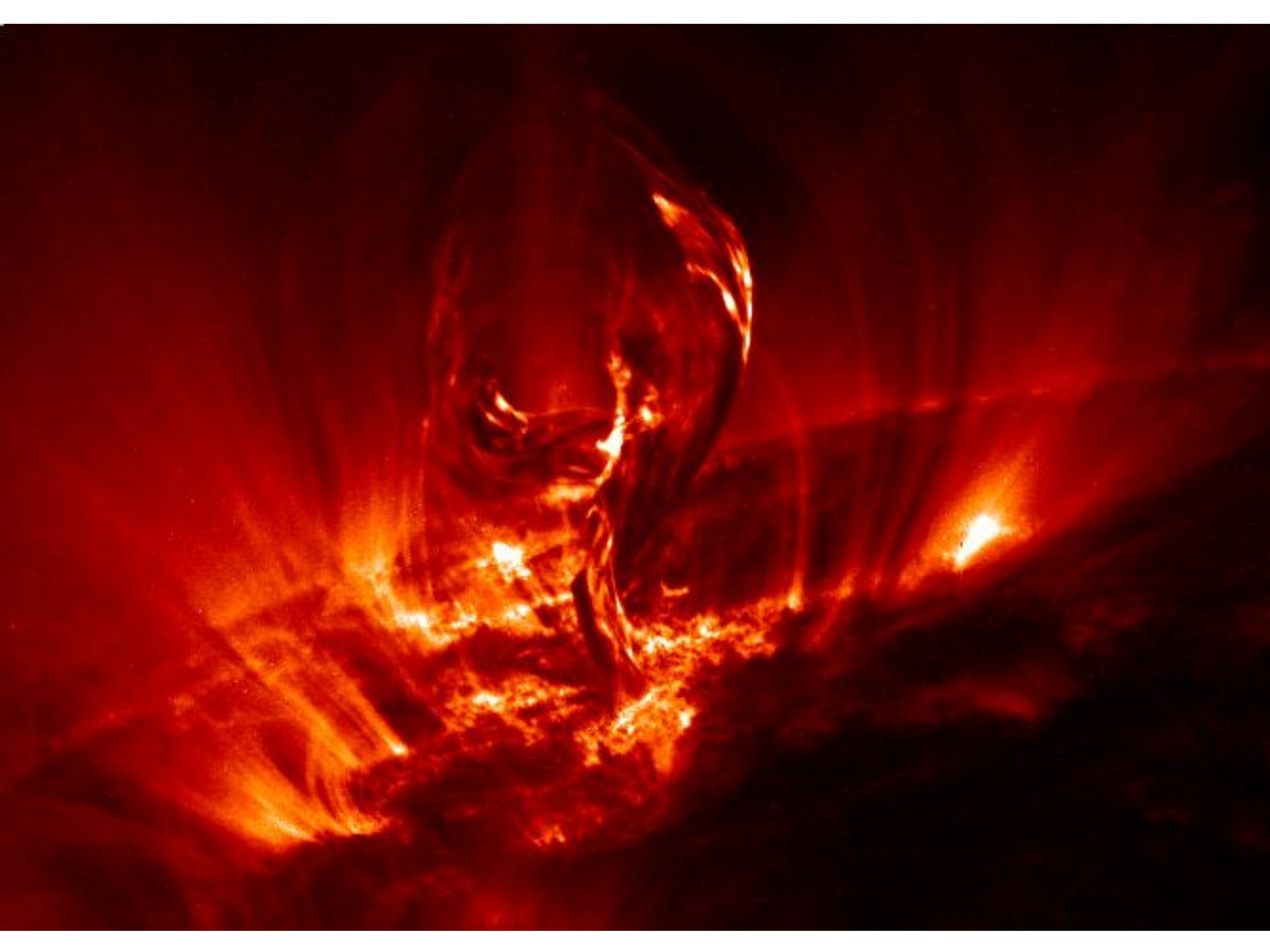
NASA caption: Giant magnetic loops dance on the sun's horizon in concert with the eruption of a solar flare—seen as a bright flash of light—in this imagery from NASA's Solar Dynamics Observatory, captured Jan. 12-13, 2015.

[Image Credit: NASA/SDO](#)

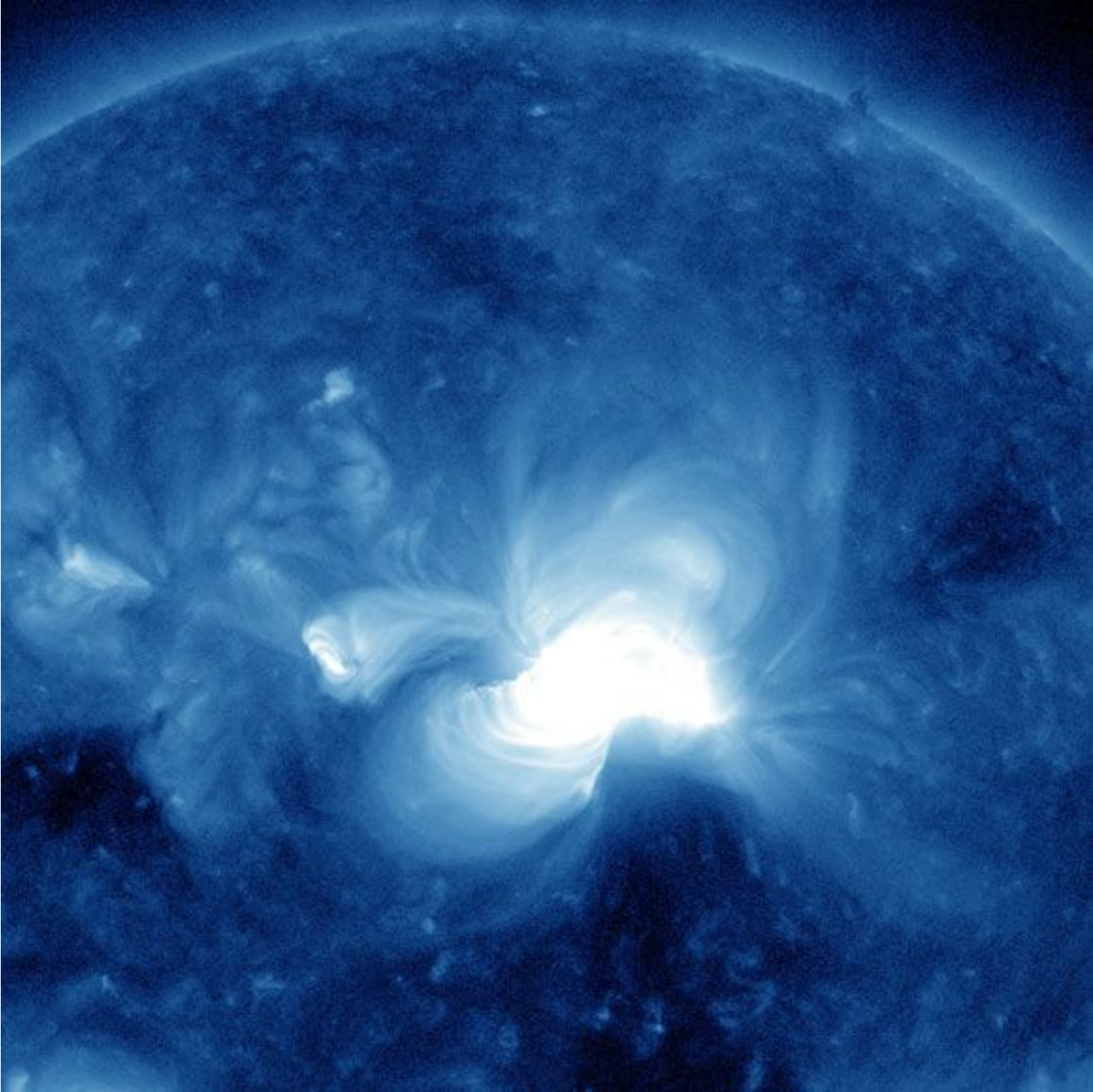










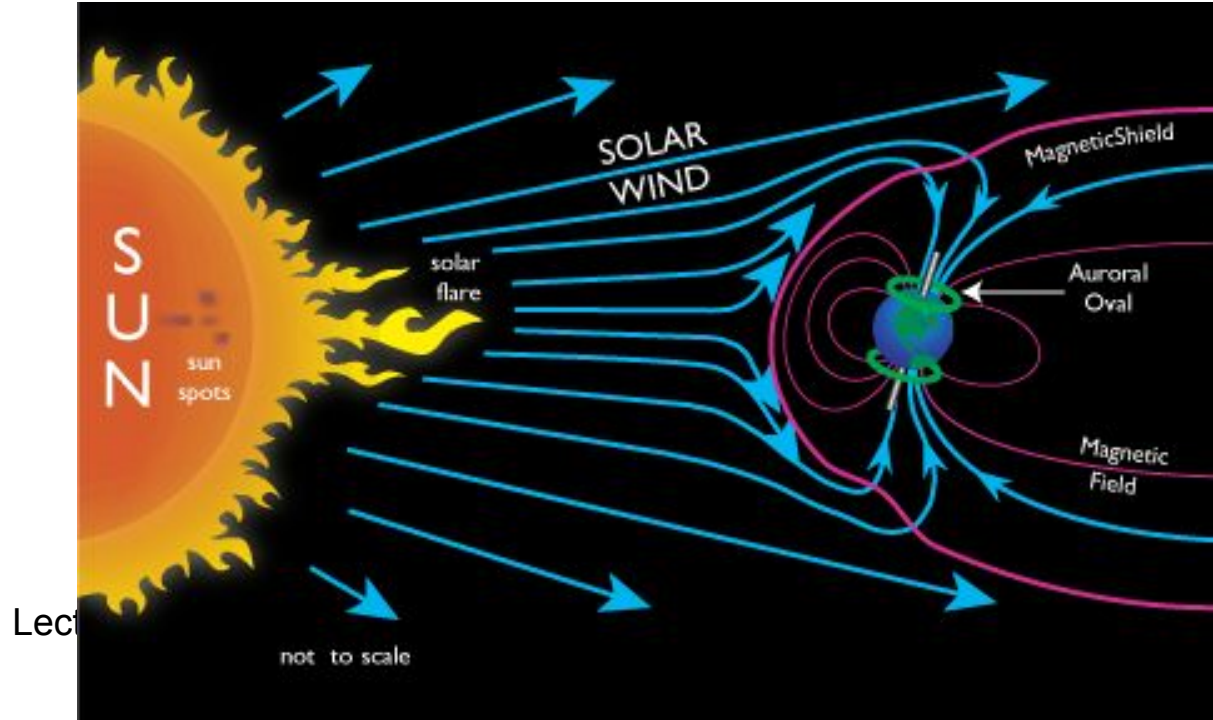


# Sun surface videos

- <https://www.youtube.com/watch?v=ipvfwPqh3V4>
- <https://www.youtube.com/watch?v=0WW1HN0iG0M>
- <https://www.youtube.com/watch?v=lpzCSZ7Eerc>
- <https://www.youtube.com/watch?v=nmDZhQAleXM>

# Solar wind

- The total number of particles carried away from the Sun by the solar wind is about  $1.3 \times 10^{36}$  per second.
- Thus, the total mass loss is about 4–6 billion tons per hour.
- Composed of:
  - electrons,
  - protons
  - alpha particles

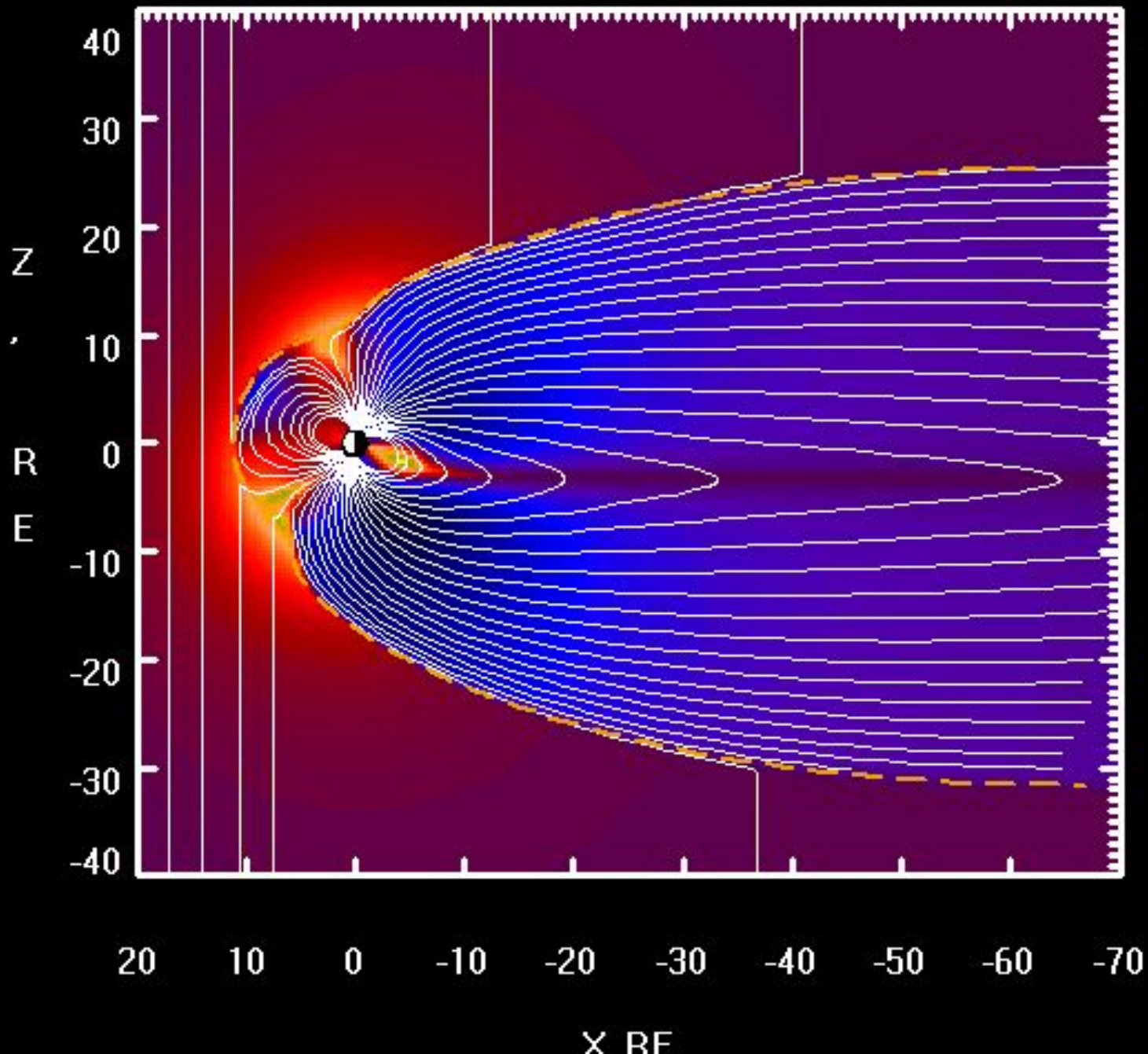


# Elementary particles flow from Sun – Solar Wind

## Aurora Borealis







# Aurora Borealis

- <https://www.youtube.com/watch?v=hsMW7zbzsUs>
- <https://www.youtube.com/watch?v=Vdb9IndsSXk>
- <https://www.youtube.com/watch?v=pjgvGiEHINs>



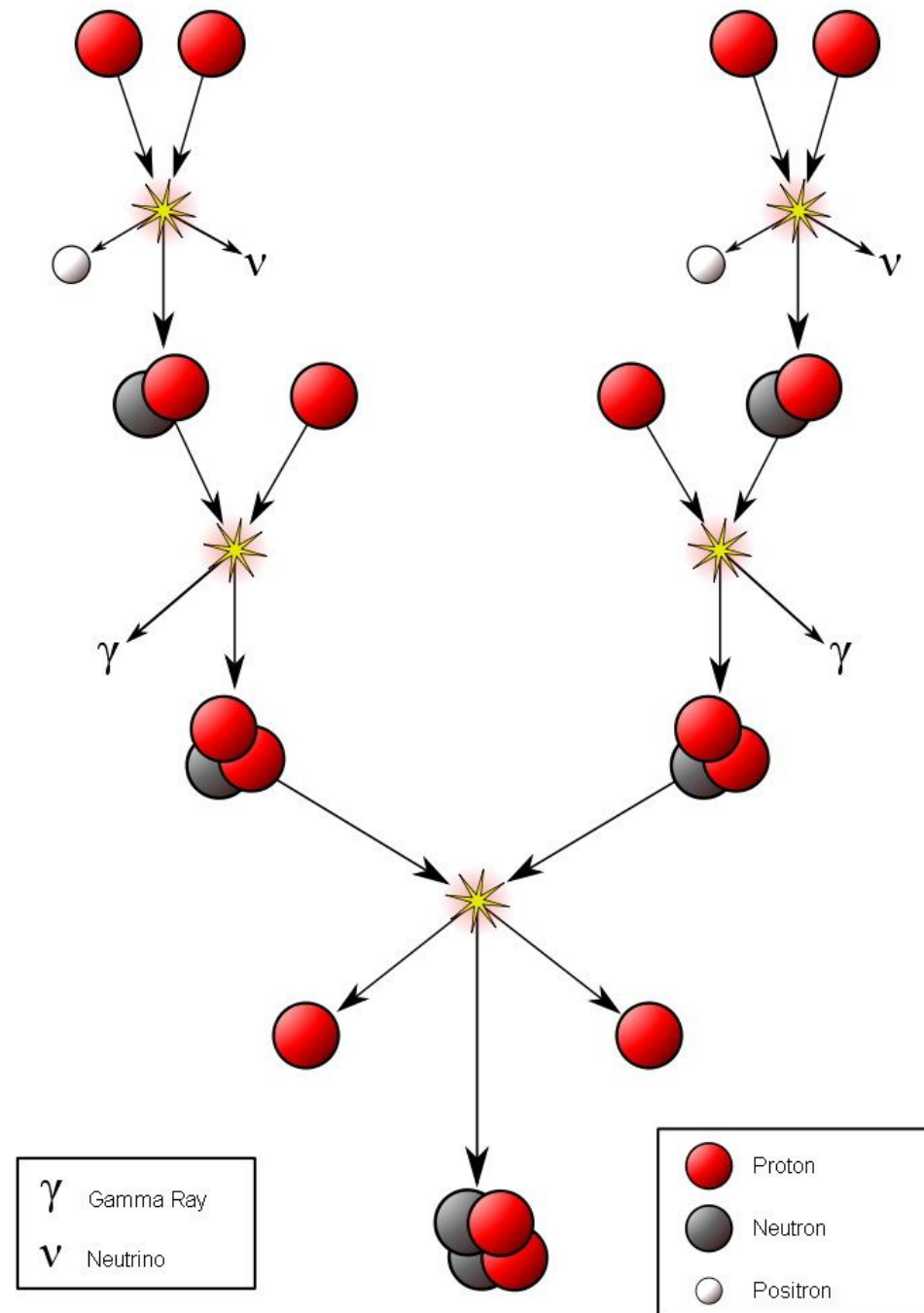
# How this energy is generated?

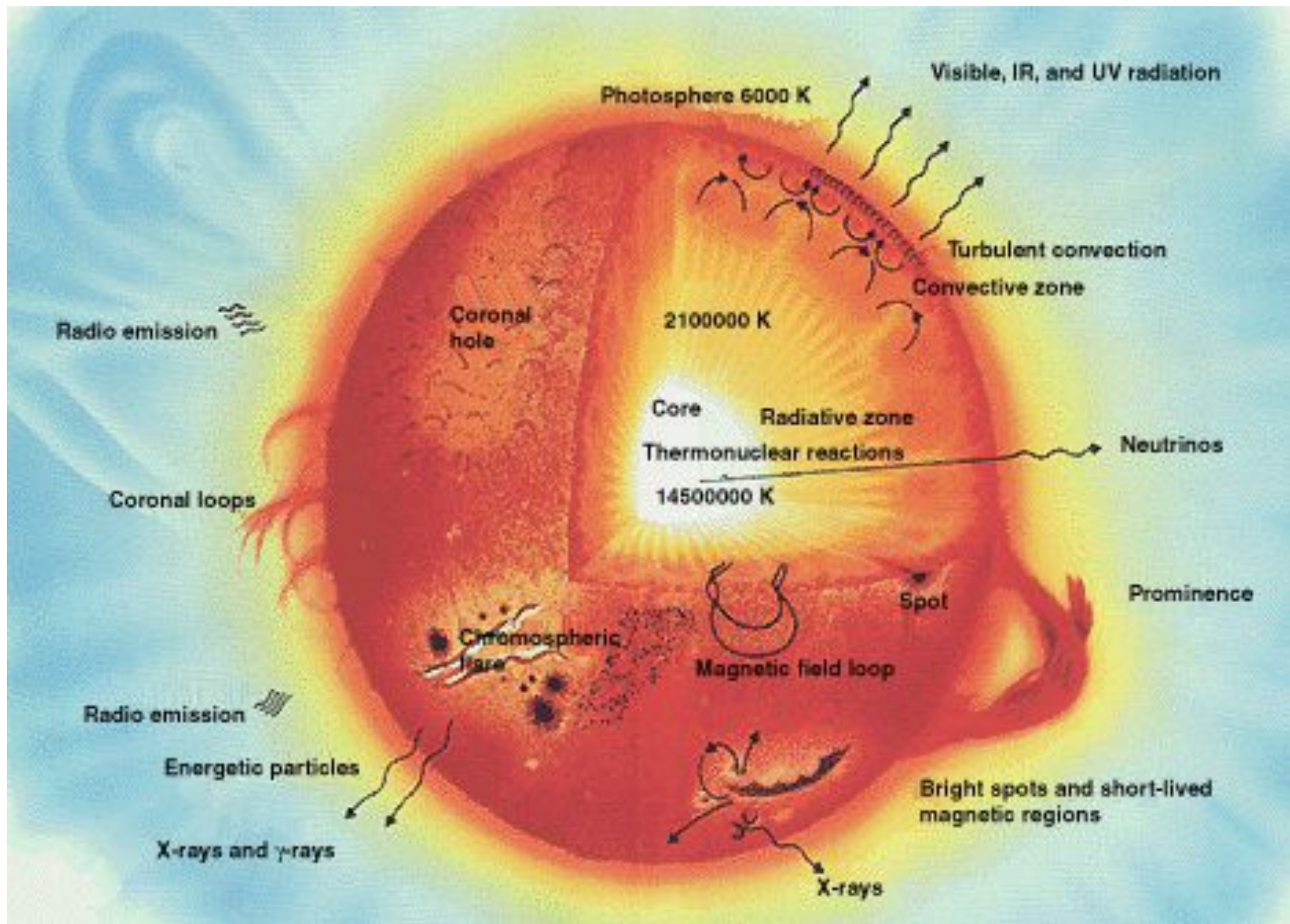
In 1920 Sir Arthur Eddington proposed that the pressures and temperatures at the core of the Sun could produce a nuclear ***fusion*** reaction that merged hydrogen into helium, resulting in a production of energy from the ***net change in mass***.

This actually corresponds to a surprisingly low rate of energy production in the Sun's core—about  $0.3 \mu\text{W}/\text{cm}^3$  (microwatts per cubic cm), or about  $6 \mu\text{W}/\text{kg}$  of matter.

For comparison, the human body produces heat at approximately the rate  $1.2 \text{ W}/\text{kg}$ , roughly a million times greater per unit mass.

Lecture # 4,





How this energy is generated?

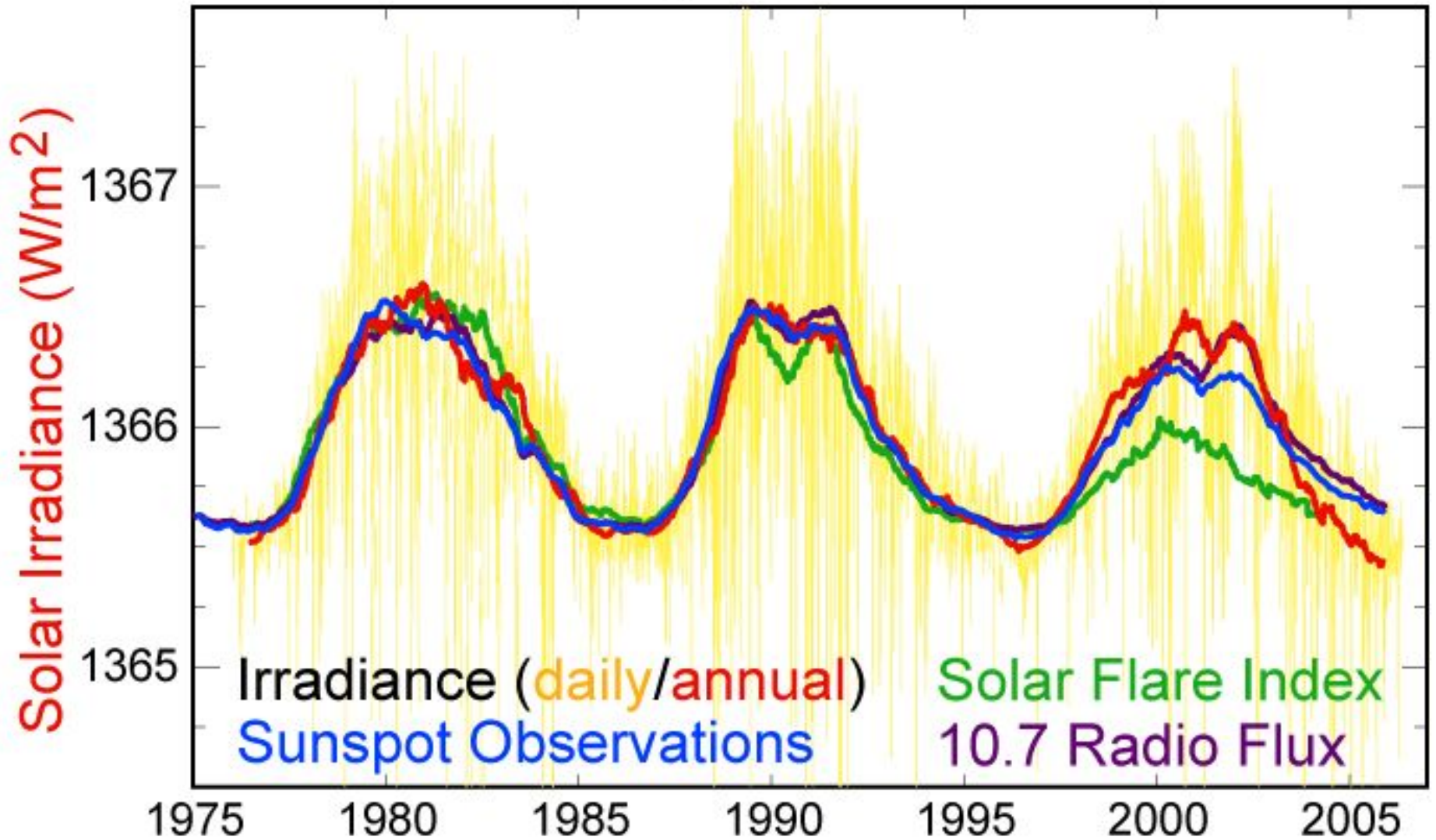
most of the elements  
in the universe had  
been created by  
nuclear reactions  
inside stars like the  
Sun.

# 1.5 The future of energy resources

- Solar Constant = 1366 W/sq.m.
- Sahara's surface area = 9,000,000 sq.km.
- If we use 10% of Sahara with 12.5% efficiency, we will get 1000 Exajoules/year!
- This is twice as much as current world consumption.
- I can see the future «Ocean Solar Power Plants», that produce Hydrogen!
- However, population grows exponentially!

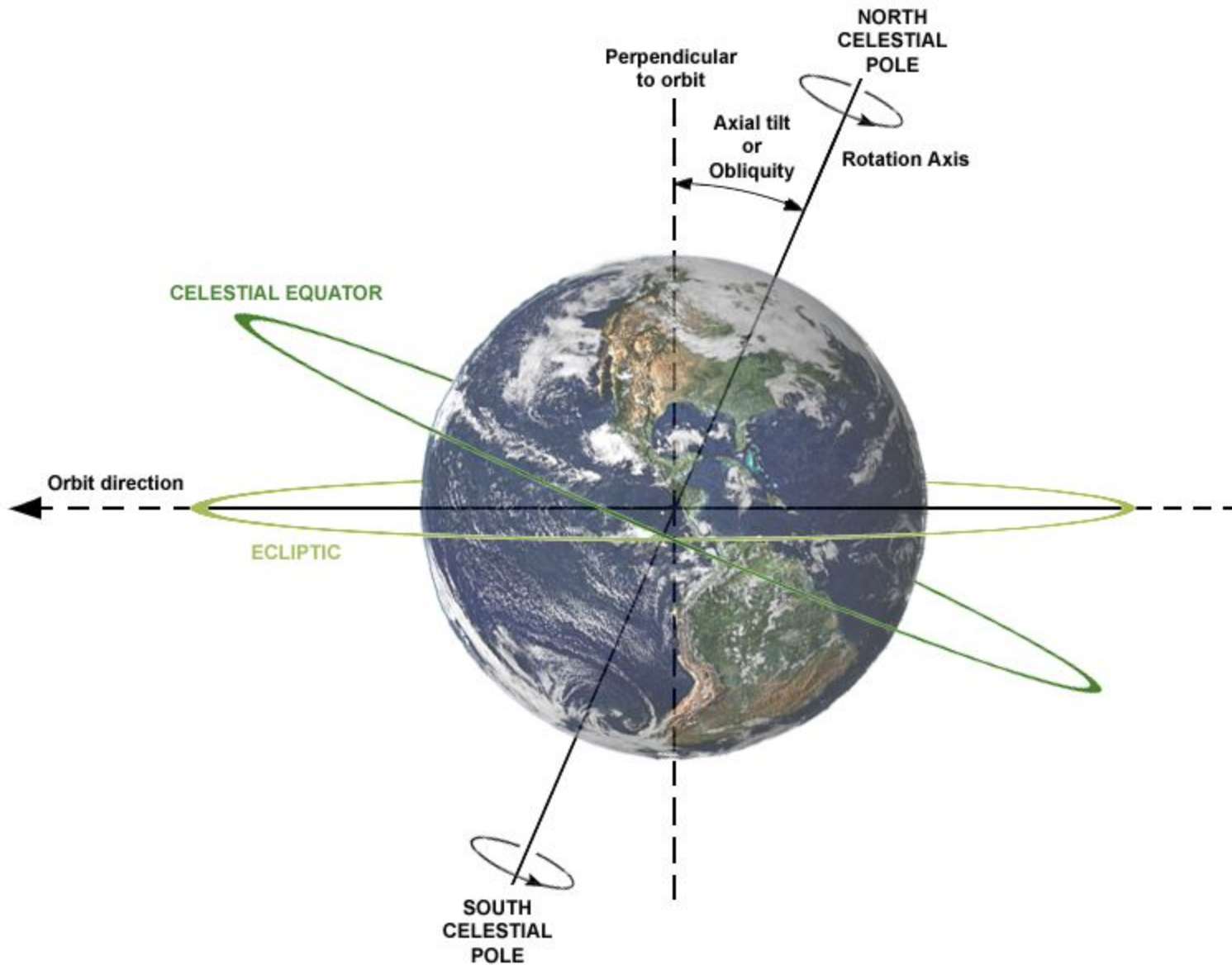


# Solar Cycle Variations



# Earth's rotation

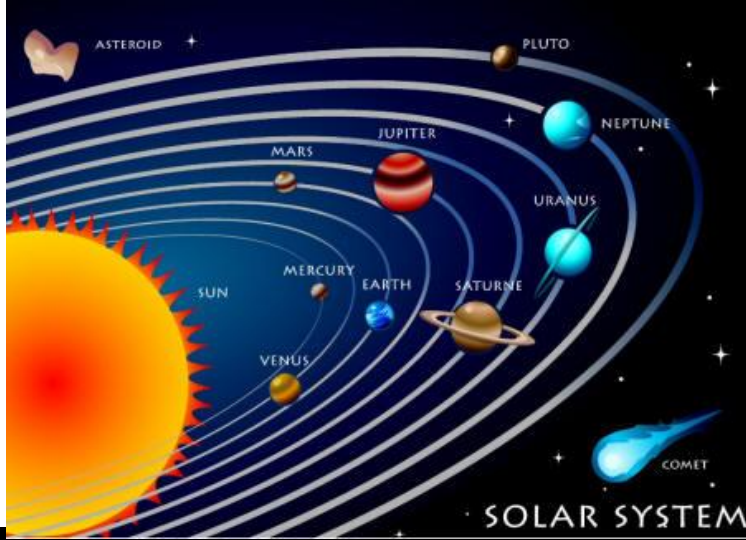
- Earth's rotation tilts about 23.5 degrees on its pole-to-pole axis, relative to the plane of Earth's solar system orbit around our sun.
- As the Earth orbits the sun, this creates the 47-degree peak solar altitude angle difference, and the hemisphere-specific difference between summer and winter.





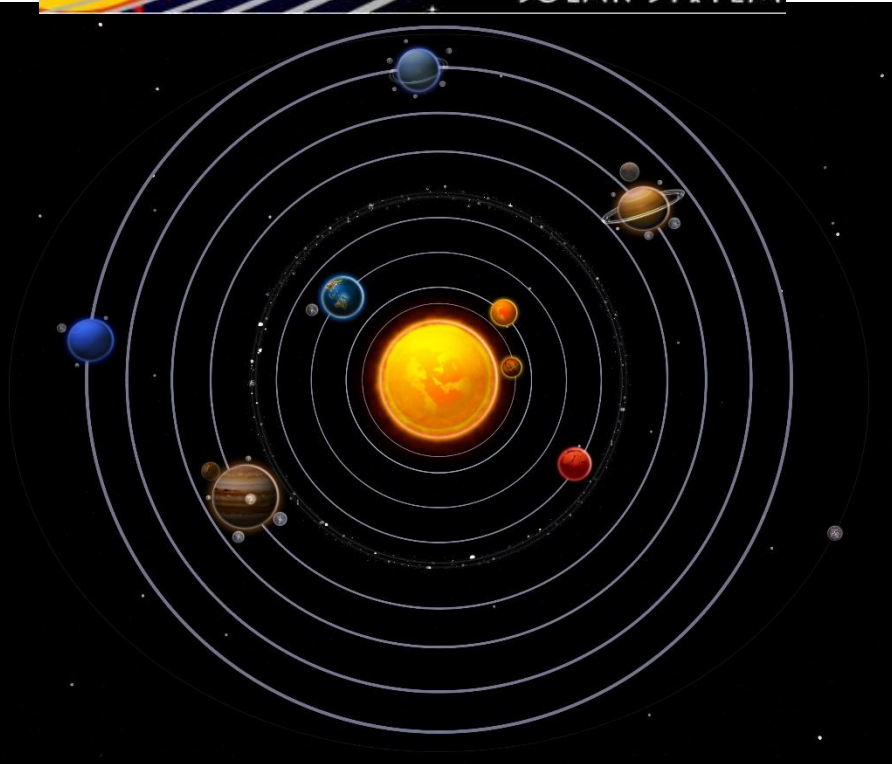
# Solar Constant

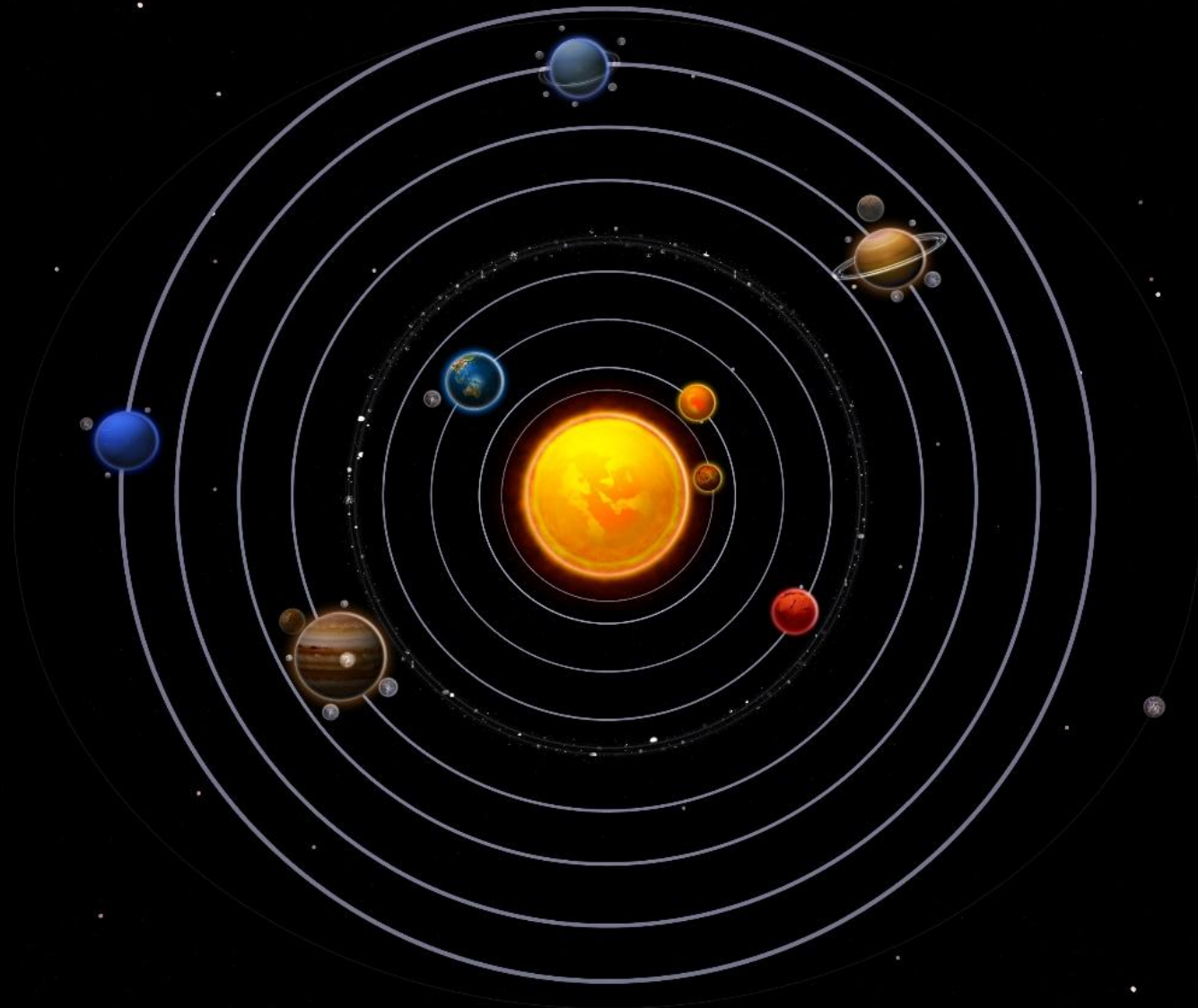
Planets, Distances and Incidences



Planet	Mean Radius (AU)	Solar Radiation Incidence (W/m <sup>2</sup> )
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Mercury	0.387	9,121
Venus	0.720	2,635
<b>Earth/Moon</b>	<b>1.000</b>	<b>1,366</b>
Mars	1.520	591
Asteroid Belt	2.500	219
3.5 AU Distance	3.500	112
Jupiter	5.190	51
7 AU Distance	7.000	28
Saturn	9.510	15
Uranus	19.000	4
Neptune	30.000	1.5
Pluto	39.480	0.9





Now: go to the article

<http://www.wired.com/2015/07/pluto-new-horizons-2/>

Lecture # 4, Solar Astronomy

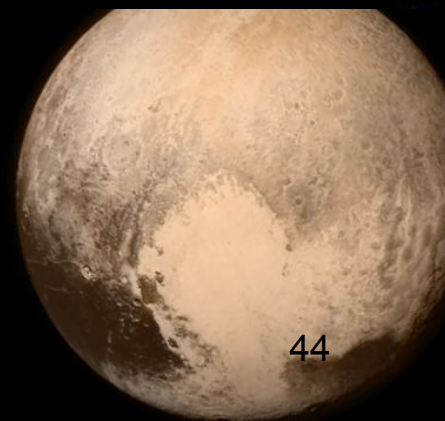
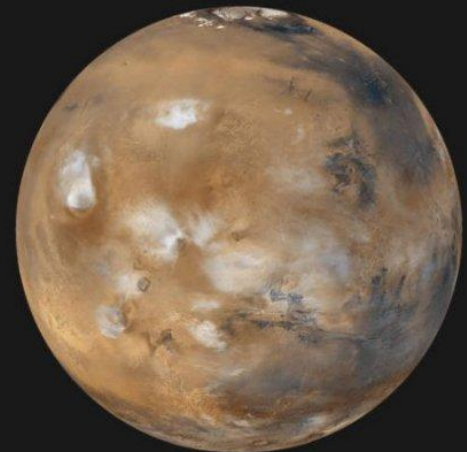
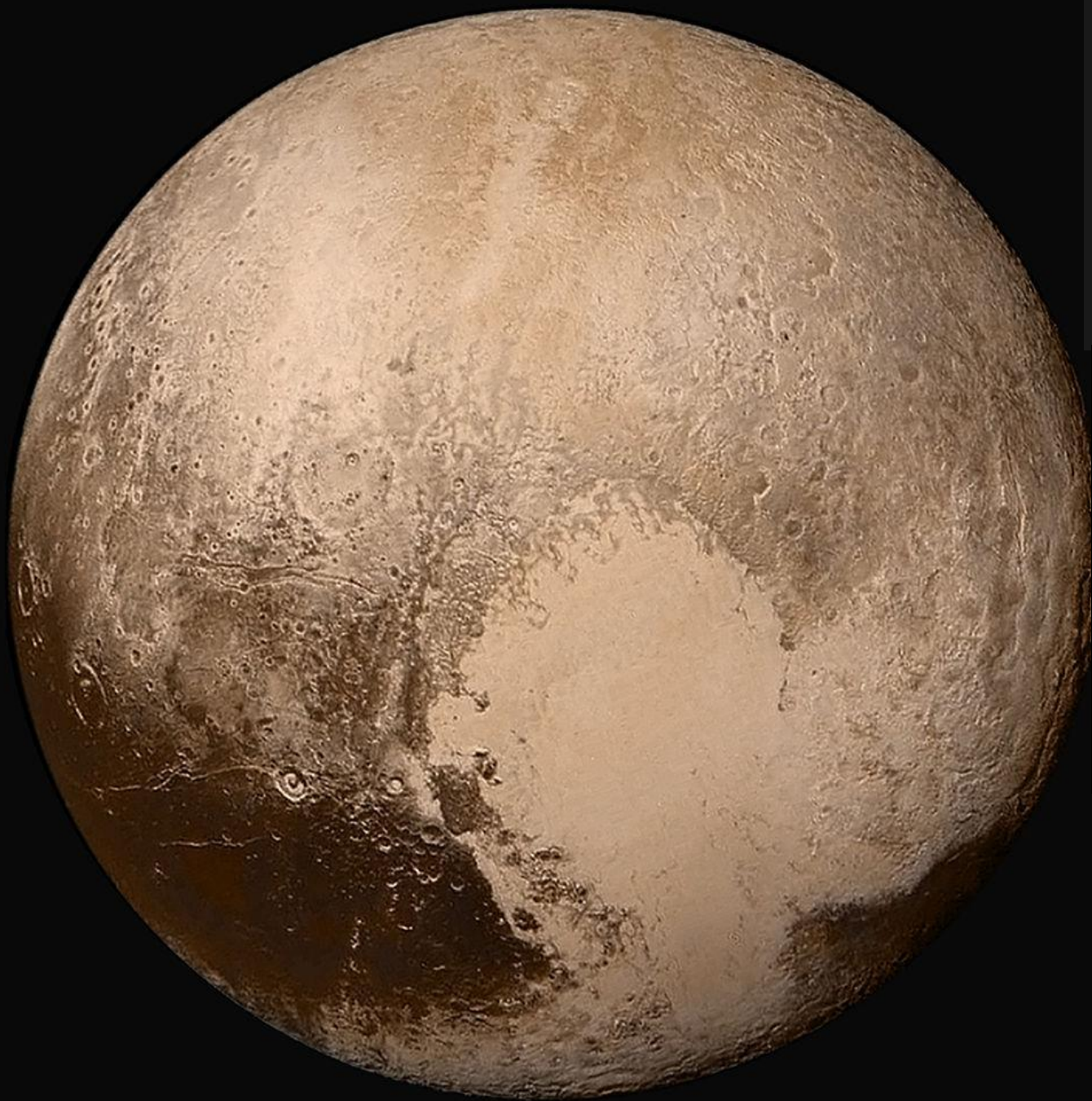


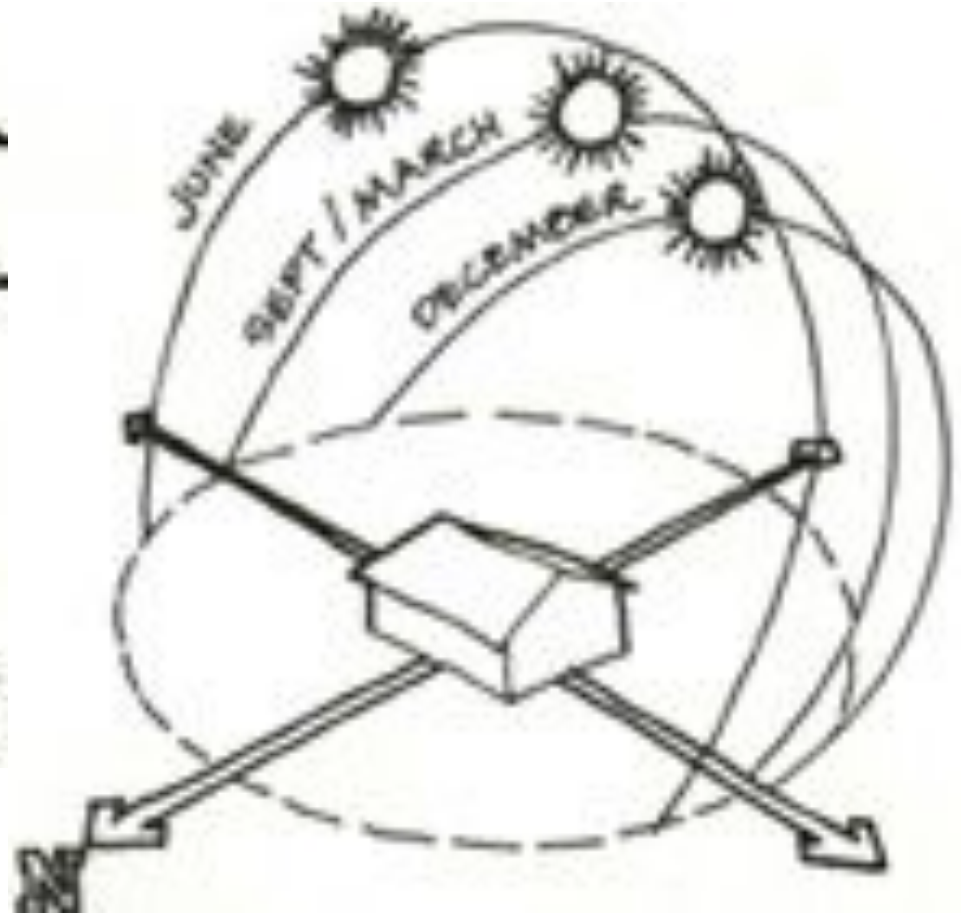
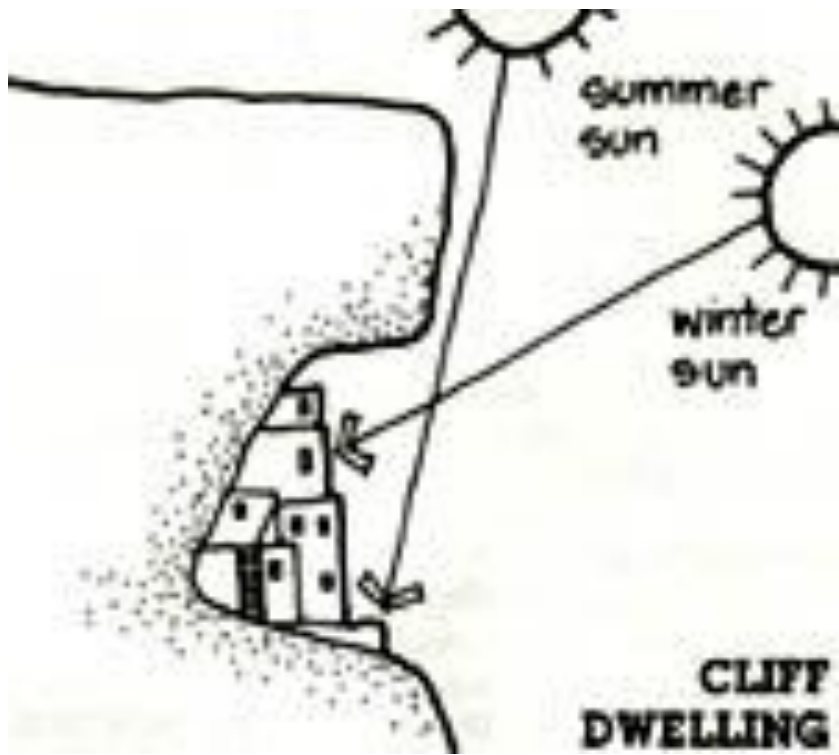
A detailed illustration of the New Horizons spacecraft in space. The spacecraft is gold-colored with a large white parabolic dish antenna. It is positioned in the foreground, angled towards the right. In the background, a large, detailed Earth with visible continents and oceans dominates the center. To the upper left, a smaller, grey Moon is visible. The background is a dark, star-filled space.

**NEW HORIZONS'  
LONG, DARK, AMAZING  
JOURNEY TO PLUTO...  
AND BEYOND  
BY SARAH ZHANG**

Lecture # 4, Solar Astronomy







# Airmass

- In astronomy, airmass is the optical path length through Earth's atmosphere for light from a celestial source.
- As it passes through the atmosphere, light is attenuated by scattering and absorption; the more atmosphere through which it passes, the greater the attenuation.
- Consequently, celestial bodies at the horizon appear less bright than when at the zenith.



# Earth Atmosphere



# Rayleigh scattering

$$\sigma_s = \frac{2\pi^5 d^6}{3 \lambda^4} \left( \frac{n^2 - 1}{n^2 + 2} \right)^2$$





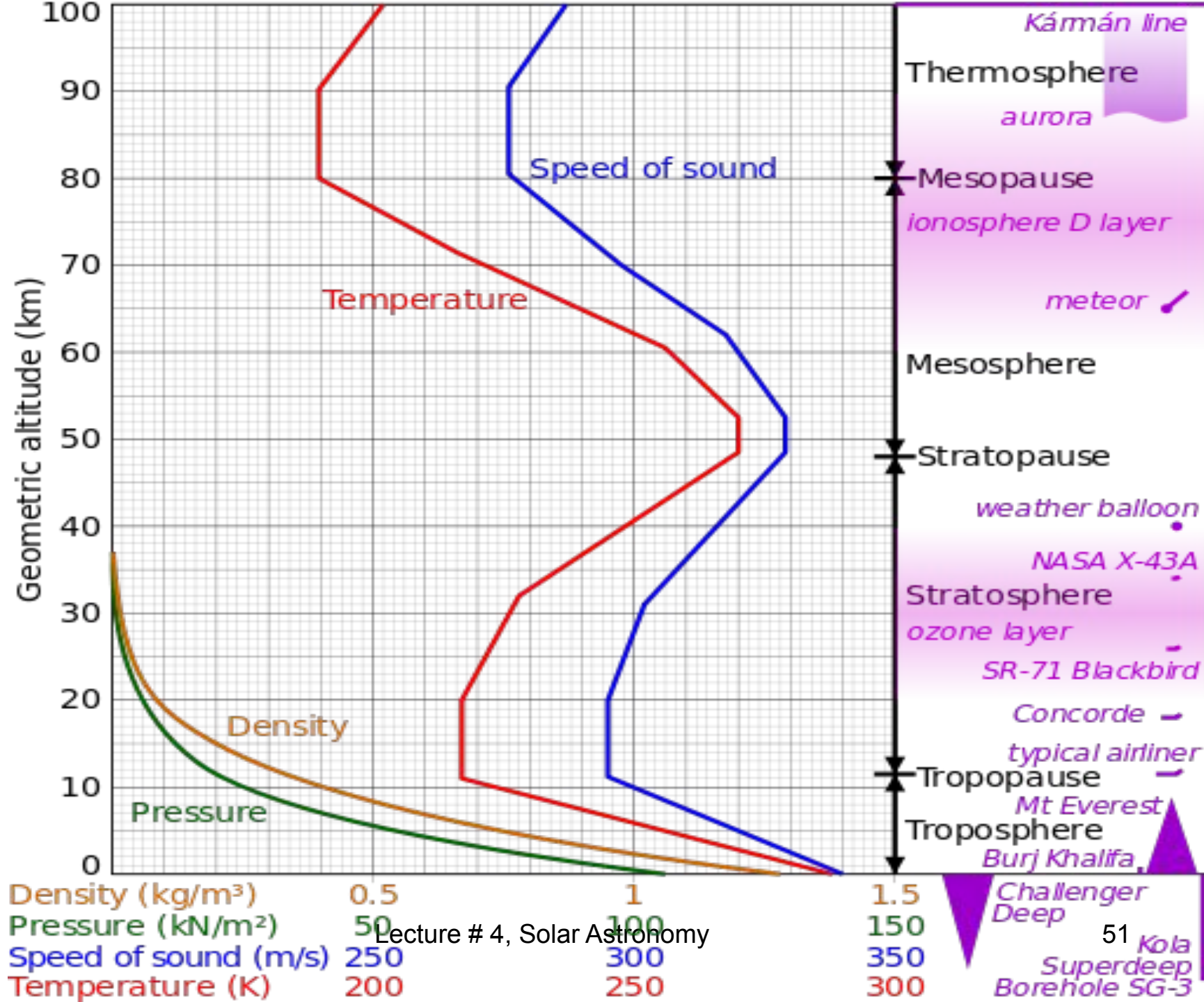
# Airmass

- “Airmass” normally indicates relative airmass, the path length relative to that at the zenith at sea level, so by definition, the sea-level airmass when the sun is at the zenith is 1.
- Airmass increases as the angle between the source and the zenith increases, reaching a value of approximately 38 at the horizon.
- Airmass can be less than one at an elevation greater than sea level.

# Airmass

- Atmosphere height = 8.5 ÷ 11 km.
- Earth's mean radius is 6371 km.
- Airmass abbreviation: AM##.
- E.g. at angle of approximately 60 degrees over horizon we have AM2, = 62% of solar constant.
- The solar panels are often rated at AM1.5
- The maximum airmass at horizon is:  
AM35.5 ÷ AM39
- At sea level, AM1 attenuates @ 27%.
- At AM10 we have 23X attenuation
- At AM20 we have >10000X attenuation

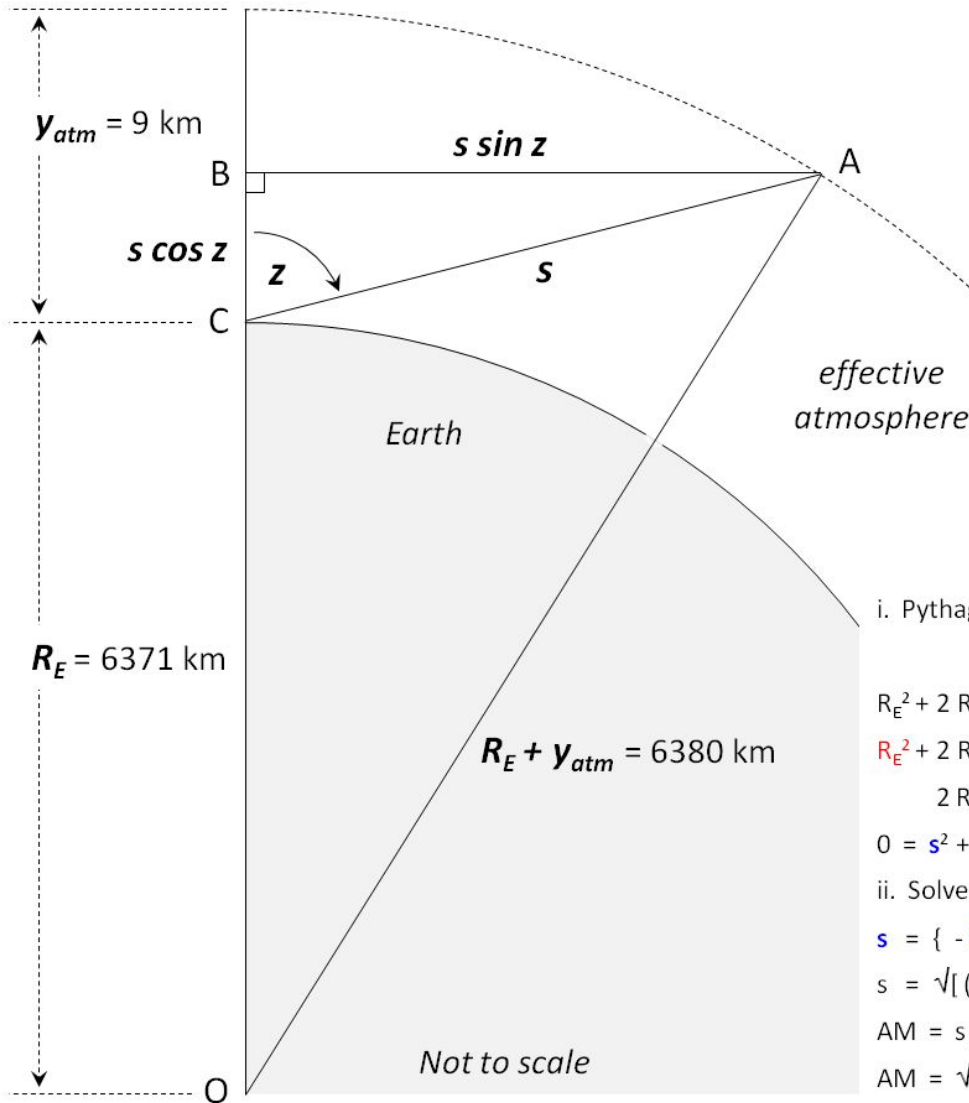
# Earth Atmosphere



# Numbers to remember

- Solar constant =  $1366\text{W/m}^2$
- Attenuation at AM1 = 27%
- Scattered light capacity  
=  $1366\text{W/m}^2 \times 27\% = 369\text{W/m}^2$
- Intensity at AM1 =  $1366\text{W/m}^2 - 369\text{W/m}^2$   
=  $997\text{W/m}^2 \approx 1000\text{W/m}^2$
- Reference Intensity =  $1000\text{W/m}^2$

# Air mass calculations



i. Pythagoras applied to right-angle triangle OAB :

$$(R_E + y_{atm})^2 = (R_E + s \cos z)^2 + (s \sin z)^2$$

$$R_E^2 + 2 R_E y_{atm} + y_{atm}^2 = R_E^2 + 2 R_E s \cos z + s^2 \cos^2 z + s^2 \sin^2 z$$

$$R_E^2 + 2 R_E y_{atm} + y_{atm}^2 = R_E^2 + 2 R_E s \cos z + s^2 (\cos^2 z + \sin^2 z)$$

$$2 R_E y_{atm} + y_{atm}^2 = 2 R_E s \cos z + s^2$$

$$0 = s^2 + 2 R_E s \cos z - (2 R_E y_{atm} + y_{atm}^2)$$

ii. Solve quadratic for  $s$  :

$$s = \{ -2 R_E \cos z \pm \sqrt{[(2 R_E \cos z)^2 + 4(2 R_E y_{atm} + y_{atm}^2)]} \} / 2$$

$$s = \sqrt{[(R_E \cos z)^2 + 2 R_E y_{atm} + y_{atm}^2]} - R_E \cos z$$

$$AM = s / y_{atm} ; r = R_E / y_{atm} \approx 708$$

$$AM = \sqrt{[r^2 \cos^2 z + 2 r + 1]} - r \cos z$$

# Notion of the Cost per peak watt installed

- “Peak Watt” = 1000W = 1kW
- Is the power produced at normal incidence of solar radiation @ 1000W/m<sup>2</sup>.
- $\$/W_p$  - Easy way to compare various solar conversion devices.
- Mostly useful for electric power generation devices, such as for: Hydro; PV; Wind, Solar Thermal Electric, etc.